

Quantum Aspects of Black Holes - Lecture 18 - 25 October 2016

Last time we discussed the naive information paradox.

The paradox is as follows.

Consider a black-hole formed from collapse



The Hawking radiation seems to emerge from the horizon which seems to have no information

about the infalling matter.
eg. we derived

$$\langle a_{\omega}^{\dagger} a_{\omega'} \rangle = \frac{e^{-\beta \omega}}{1 - e^{-\beta \omega}} \delta(\omega - \omega')$$

For the modes outside independent of
the details of the collapse.

This suggests the black hole evaporates in
a time $t_{\text{th}} \propto M^3$.

If the final radiation is thermal, then we have a contradiction. The Schrodinger equation

$$i \frac{\partial |\psi\rangle}{\partial t} = H |\psi\rangle$$

does not allow a pure state to turn into a thermal state.

But this is not a paradox because

1) Pure states can mimic mixed states very closely.

2) Hawking's calculation is not precise enough to distinguish between a final pure state and final mixed state.

To understand claim 1, consider a closed system where the Hilbert space has dimension e^S .

This means that a basis for the Hilbert space is given by

$$|1\rangle \dots |e^S\rangle.$$

But a typical state is given by

$$|\psi\rangle = \sum a_i |i\rangle \quad \text{with} \quad |a_i|^2 = 1$$

so the Hilbert space can be identified with

$$\mathbb{C}P^s$$

We now want to understand observables or typical states in this Hilbert space.

This means that we pick a random pure state using a democratic probability distribution, $|\psi_{\text{typ}}\rangle$, in this Hilbert space and ask for the expectation value of an operator A

$$\langle \psi_{\text{typ}} | A | \psi_{\text{typ}} \rangle$$

This is computed by averaging $\langle \psi | A | \psi \rangle$ as $|\psi\rangle$ runs over the whole Hilbert space

$$\langle\langle A \rangle\rangle = \int_{\mathbb{C}P^n} \langle \psi | A | \psi \rangle dM_{|\psi\rangle}.$$

where $dM_{|\psi\rangle}$ is a measure factor that assigns equal weights to equal volumes on $\mathbb{C}P^n$.

Such a measure is easy to find. We simply consider

$$\frac{1}{V} da_1 \dots da_{e_s} \delta(|a_1|^2 + \dots + |a_{e_s}|^2 - 1) = d\mu$$

where V is the normalization required to ensure

$$\int d\mu = 1$$

We can work out the normalization explicitly

$$\int \frac{e^{it(\frac{1}{2}|a|^2-1)}}{2\pi} dt \frac{1}{\pi} d|a|^2 \quad (\pi) e^s e^{-\frac{1}{2}|a|^2}$$

↑
from angular
part of a_i

$$\int \left(\frac{-1}{it - \frac{1}{2}} \right) \frac{e^{-it}}{2\pi} dt \quad (\pi) e^s$$

complete t contour through the lower half plane, pick up pole at $t = -i$ with residue

$$\frac{(\pi) e^s}{\Gamma(e^s)}.$$

Now consider an operator A . We claim that

$$\langle\langle A \rangle\rangle = \frac{1}{e^S} \text{tr}(A).$$

This means that the expectation value of A in a typical state is **exactly the same** as in the **identity density matrix** $\rho = \frac{1}{e^S} I$ that assigns equal probability to all states.

This is easy to prove.

Pick a basis so $A = \sum_i A_{ii} |i\rangle\langle i|$

then

$$\langle\langle A \rangle\rangle = \int \sum_i |a_i|^2 A_{ii} \frac{\delta(\sum_i |a_i|^2)}{V} d^2 a_i$$

$$= \sum_i A_{ii} \langle |a_i|^2 \rangle.$$

Since $\langle |a_i|^2 \rangle = \alpha^2$ independent of i and since

$$\langle \sum_i |a_i|^2 \rangle = e^S \alpha^2 = 1$$

we must have

$$\alpha^2 = e^{-S}$$

$$\Rightarrow \langle\langle A \rangle\rangle = e^{-S} \sum_i A_{ii} = e^{-S} \text{Tr}(A).$$

This is a simple, but it means that if I hand you a pure state you will get the same $\langle A \rangle$ as in the mixed state.

But maybe pure states can be distinguished by fluctuations? i.e. perhaps the average over pure states is $\frac{\text{tr}(A)}{e^S}$ but individual pure

states see large fluctuations from this value?

The next theorem rules this out.

Claim: Consider some operator A . Then

$$\int \left(\langle \psi | A | \psi \rangle - \frac{1}{e^S} \text{Tr}(A) \right)^2 d\mu.$$

$$= \frac{1}{(e^S + 1)} \left[\frac{\text{Tr}(A^2)}{e^S} - \frac{(\text{Tr}(A))^2}{e^S} \right]$$

so the average deviation is suppressed
by $e^{-S/2}$!

Proof: In the same basis, we now consider

$$\begin{aligned}
 & \int d\mu \left(\sum_i A_{ii} |a_i|^2 - \frac{1}{e^s} \left(\sum_i A_{ii} \right)^2 \right) \\
 &= \sum_{i,j} \int d\mu \left(A_{ii} A_{jj} |a_i|^2 |a_j|^2 + \frac{1}{e^{2s}} \sum_i A_{ii} A_{jj} \right. \\
 &\quad \left. - 2 \frac{A_{ii} |a_i|^2 A_{jj}}{e^s} \right) \\
 &= \sum_{i,j} \int d\mu A_{ii} A_{jj} |a_i|^2 |a_j|^2 - \left(\frac{\sum_i A_{ii}}{e^s} \right)^2
 \end{aligned}$$

$$\int \pi \frac{d|a_i|^2}{V} e^{-\sum |a_i|^2} e^{t(\sum |a_i|^2 - 1)} |a_i|^4 = 2 \frac{\int \frac{e^{-it}}{(it - \varepsilon)^{s+2}}}{\int \frac{e^{-it}}{(it - \varepsilon)^s}}$$

$$= \frac{2}{e^s (e^s + 1)}$$

and $\int \pi \frac{d|a_i|^2}{V} e^{-\sum |a_i|^2} e^{t(\sum |a_i|^2 - 1)} |a_i|^2 |a_j|^2 = \frac{1}{e^s (e^s + 1)}$

for $i \neq j$

So the original integral becomes

$$\sum_{i,j} A_{ii} A_{jj} \left(\frac{2 \delta_{ij}}{e^s(e^s+1)} + \frac{(1-\delta_{ij})}{e^s(e^s+1)} \right) - \left(\frac{\sum A_{ii}}{e^s} \right)^2$$

$$= \sum_i \frac{A_{ii}^2}{e^s(e^s+1)} - \left(\sum A_{ii} \right)^2 \left(\frac{1}{e^s(e^s+1)} - \frac{1}{e^{2s}} \right)$$

$$= \frac{\text{tr}(A^2)}{e^s(e^s+1)} - \frac{\text{tr}(A)^3}{e^{2s}(e^s+1)} \quad \text{as claimed!}$$

This is a general result valid for a statistical system.

If the system is in a pure state with energy

$$E \in [E_0 - \Delta E, E_0 + \Delta E]$$

then there are e^S states in this interval.

A typical linear combination of energy eigenstates in this interval is indistinguishable from the microcanonical mixed state.

Now we return to Hawking radiation.

We have shown that 2-pt functions have a "thermal" form. But this is misleading because pure-states look exponentially close to thermal states!

Hawking's calculation is not precise enough to lead to a paradox.

Can we improve Hawking's computation to obtain a paradox. To correct the computation we need to take the back-reaction of the radiation on the geometry into account.

the energy of typical Hawking quanta is

$$E = \frac{M_{\text{Pl}}^2}{M}$$

Recall the gravitational coupling constant is

$$\frac{E^2}{M_{Pl}^2} = \frac{M_{Pl}^2}{M^2} \text{ in this case.}$$

But $\frac{M_{Pl}^2}{M^2} = \frac{1}{S}!$

So by correcting the computation perturbatively
we can develop a power series in $\frac{1}{S}$

But the difference between pure and mixed states is **non-perturbative** in this coupling.

$$(\hat{T} e^{-S} = e^{-1/\text{coupling}})$$

And Q.F.T in curved spacetime cannot deal with such effects.

To refine the information paradox, we need other arguments!