31 March 2021 Lecture 21: Paradoxes for large Ads black boles

We now consider large black holes in global Ads.

For such black holes

 $ds^2 \longrightarrow -F(r)dt^2 + dr^2 F(r) + r^2 dt^2_{d-1}$

where $F(r) = 1 - \frac{m}{\sqrt{d-2}} + r^2$ $\sum M = 8\pi^{2-\frac{3}{2}} \Gamma(d) GM$

The picture we have in mind is as sollows

R region of interest tinfalling matter

We will later consider eternal black holes but for now, we are thinking of pure states.

Difference vetween large Ads black holes and evaporating black holes

Large Ads black holes dominate the microcanonical ensemble.

IF one takes a typical state at high energies, in a theory in Ads, it corresponds to a large t.t.

et most states are large black holes atypical states

evaporating lh are just the opposite. the entropy of Hawking radiation rentropy of U.h. (S)

So if pick a state "at random" with prob 1-es-s

it will be a gas of Hawking radiation.

-most states are gas of gravitons lother particles eraporating black holes.

large l.h. has positive specific heat

evaporating U.L. has negative specific heat

This will be important below as we will take a trace over the microcanonical ensemble to obtain our paradoxes.

The main Eakeaway is the following. The following two statements are inconsistent:

1) Typical pure states at high-enough energies are black lots with empty interiors where one can use FFT

2) Dot in the interior are described by the same operators in the OFT For different microstates. T

EExplain. May seem like an innocuous assumption but will turn out to le important?

Let us recall what effective field theory Lells us

Consider a lulk scalar field propagating in the l.h. geometry

By smearing the field appropriately, we can extract

Modes Que by smeaning the Field outside the O horizon

2) Modes Q by Smeaning the Field inside the horizon

[Emphasize Ehat Ehis Follows From a 2-pt Fn across the horizon.]

 $\langle \psi \rangle \widetilde{a}_{\omega} \widetilde{a}_{\omega}^{+} | \psi \rangle = \langle \psi \rangle a_{\omega} a_{\omega}^{+} | \psi \rangle = \frac{1}{1 - \beta^{-}} \beta \omega$ $\sum aw, aw^{+} \sum = \Delta$ $\Delta \omega = 1$ $\chi \gamma Q \qquad \chi \gamma = e^{-\beta w/2}$ $\chi - e^{-\beta w}$ Law, aw Z = 0

We must have

Furthermore

IH, QuJ= -WQW

 $\Sigma H = Z = W Q W$

This is important so let us briefly understand how this is derived.

Megative occupancy paradox we stated that in typical states and $\SigmaH, \widetilde{aw}S = w\widetilde{aw}$ We also know [purely based on kinematic] that expectation values in a typical state are close to thermal expectation values. Ethis also uses equivalence of ensembles, and we are not writing therms here.]

So, we need $T = 1 + Ex(e^{-BH}a_{w}a_{w}^{+}) = 1$ $Z(B) + (e^{-BH}a_{w}a_{w}^{+}) = 1$ Now, using the cyclicity of the trace

Using the H, and ommutator



50 C-PWT-1 $\overline{\}$ $-\frac{e}{1-e^{-BW}}$ 50 $\overline{)}$ But T is the expectation value of a positive operator! so this negative result is absurd!

What does the paradox imply?

1) It suggests that if we look for fixed a that behave correctly in typical states, we can into a paradox. EDDES not rule out do that behave correctly in a small fraction of states] 2) It was used to argue for

a) either large black holes in Ads have a firewall behind their porizon

1) The boundary theory does not describe the interior of typical states and must be ypical sypplemented with additional dof

3) Note that holography of information) complementarity is **insufficient** to resolve this.

We made no assumption about the Factorization of the Hilbert space

ut this paradox is inapplicable (as it stands) to black holes in Flat space or to small black holes in Ads.

These never dominate the ensemble!

[Emphasize again, by reviewing the trace argument]

So it is possible this paradox is a problem in Ads and not for more realistic black holes.

we expect Err states with a smooth horizon.

Using $\left(\frac{\partial}{\partial w} - e^{-\frac{\omega}{2}}\right) \left(\frac{1}{2}\right) = 0 = \left(\frac{\partial w}{\partial w} - e^{-\frac{\omega}{2}}\right) \left(\frac{1}{2}\right)$ for states with a smooth horizon, we expect.

 $M_{a} = \frac{1}{1 - e^{-Rw}} \left[\left(a_{w}^{+} - e^{-Rw/2} a_{w} \right) \left(a_{w}^{-} - e^{-Rw/2} a_{w} \right) \right] + \left(a_{w}^{+} - e^{-Rw/2} a_{w} \right) \left(a_{w}^{-} - e^{-Rw/2} a_{w} \right) \right]$

consider the operator

The infalling number operator

Now consider eigenstates of the Schwarzschild number operator



 $N_{u}(n) = O(n)$

Notice that.

 $\langle n \rangle q_{\omega} q_{\omega} \rangle \langle n \rangle = 0$

This is because $N_{\omega} a_{\omega} | m = n a_{\omega} | m = d = d$ LUE

 $N_{M} q_{+}^{+} | m \rangle = (m + m) q_{+}^{+} | m \rangle$

Plso $2n a_w a_w^+ \ln 7 > 1$ $2n a_w a_w^+ \ln 7 > 1$ $< n | N_{a} | n > = < n | + (a_{w}^{+} - e^{-\beta w | 2} a_{w})(a_{w} - e^{-\beta w | 2} a_{w}) | n >$ $2e^{-\beta \omega}$ now note that But [14, NW] = 0 [neglecting]

Now consider a trace over the micro-ens. EE Eot D <YINA/47=1 SKEINA/EY = 0 OS EE EO-D But since [Nw, H] =0, we can perform a change of basis in the trace and evaluate the trace in the number eigenbasis. But 1 & KNINAIN> > 2e-BW es [-e-BW

So we again have a contradiction.

The same comments as previously apply here.

This paradox involves a trace. It does not preclude the possibility that a small Fraction of energy eigenstates have <Na?=0.

We will later see how these paradoxes can be resolved via state-dependence.

But, we now want to emphasize an important point.

9E is sometimes believed that these paradoxes afflict only single-sided black holes.

We now explore how the same paradoxes also apply to the eternal black hole.

Review of the Eternal black hole

we consider the maximal extension of the geometry

 $dS = -F(x) dE^2 + dx^2 + x^2 dE_{-1}$

 $F(r) = 1 - \frac{m}{r^{d-2}} + r^2$

Note there is no "E-Ja" in the metric above.

geometry looks like this The right future horizon R «right asympt. region 1est asymptotic region Kright past horizon believed that this is dual to It is the thermofield doubled state of EWO CETS $\frac{1}{2(B)} = \frac{-\beta E(2)}{E}$ 17> =

where the sum runs over all eigenstates I type in current version of the review.] sum is missing!

This is the canonical example of

ER = EPR

Note a remarkable feature of this proposal

The two CFTs are entangled but noninteracting.

In general, when this happens no unitary on one system can affect anything on the other system.

But here we Find something interesting worldline of observer ni zgmujotu From the right N'(F) Stack-mare created unitary action on the left.

The experience of a right infalling observer can le affected ly a shockwave thrown From the left!

Note that this only happens in the interior we cannot send any signals From the left boundary to the right one.

We now point out that the following 3 assumptions are inconsistent

n Eternal vlack hole is dual to the thermofield double state

21 Dof in the interior are described by the same operators in the eternal b.h. and a class of states related to the eternal which by Hamiltonian time evolution

3) Disentangled states IE, E> are not connected by a wormhole.

Explain last point.